

# Retrievals in Presence of Sunlight: The Future is Looking Bright

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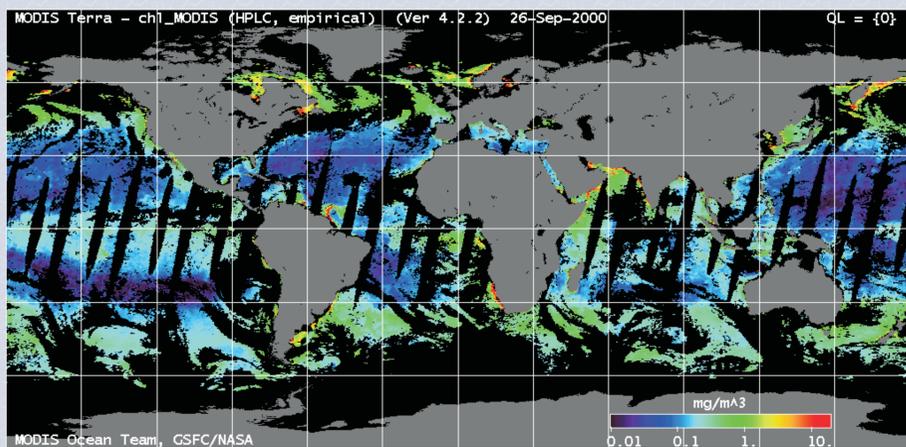
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## Sunlight: Can we do better?

- Our ability to retrieve information from the ocean and the overlying atmosphere depends on how well we can describe the ocean BRDF in our forward models.
- Sunlight only looks smooth when we average the fine structure over time and space, or look at it from great distances.
- Accurate description of ocean BRDF is becoming as important as aerosol treatment in coupled atmosphere-ocean RT models for remote sensing retrievals.
- An accurate description of sunlight together with a realistic radiative transfer model can reduce the area currently masked away as unusable data.



Above: Chlorophyll, 26 Sep 2000; alternate black stripes are missing data due to glint (others are gaps between swaths)

## Sunlight: is it just contamination?

Sunlight is like having a bright lamp at the surface

- atmospheric spectral transmission measurements are easiest to interpret

We could use sunlight regions to retrieve:

- water vapor column (Kleidman..., 2001)
- aerosol absorption (Kaufman..., 2002)

Prerequisite: accurate description of polarized sunlight at the bottom of the atmosphere

- CO<sub>2</sub> column amount at 1.5/2.2 μm wavelengths



Left: Sunlight as seen from far away. Note how a significant portion of the earth's disk is in the glint region. Extra solar planets with large water bodies might one day be detected from their glint.



Right: Sunlight seen from ground level.

Below: Sunlight as seen from the shuttle at 800 km.



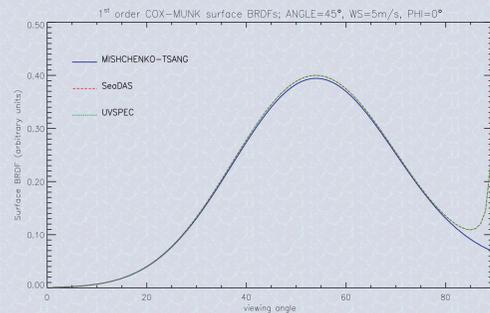
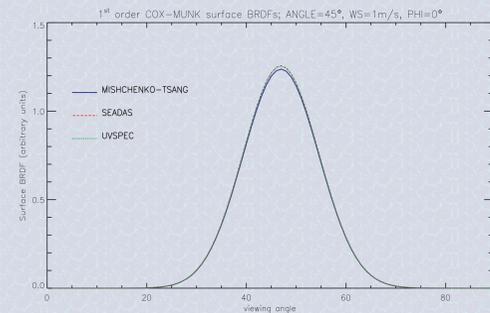
Below: Sunlight around the Hawaiian Islands as seen from 800 km up. Note the glint variation due to wind/wave conditions.



## Modeling Sunlight

Sunlight models are usually based on the work of Cox and Munk (1954!). Techniques used:

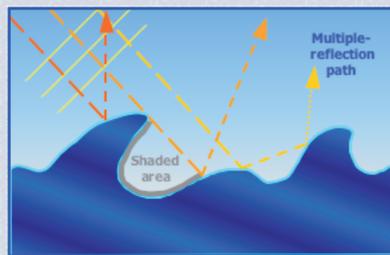
- surface statistics combined with an electromagnetic boundary value formulation
- surface realizations (based on statistics) combined with Monte Carlo calculations



## Comparisons of sunlight models

Models tested so far show small differences in

- ~ peak radiance in the specular direction
- ~ the angular direction of maximum specular reflection
- ~ the width (FWHM) of the glint region
- ~ treatment of shadowing effects
- ~ treatment of multiple reflection paths
- ~ inclusion of polarization, wind direction



Above: Surface BRDF for a sea-surface calculated with three different models. All models use the Cox/Munk wave-slope distribution. Notable differences are only apparent for high viewing angles.

## Correcting for Sunlight

Our goal is to provide algorithms that accurately will determine if glint is present, its magnitude, and polarization. This information can further

- be incorporated in forward models, enabling retrievals further into the glint contaminated region
- enable retrievals of atmospheric parameters even when the intensity of the glint is too great for accurate determination of the radiance emanating from inside the water

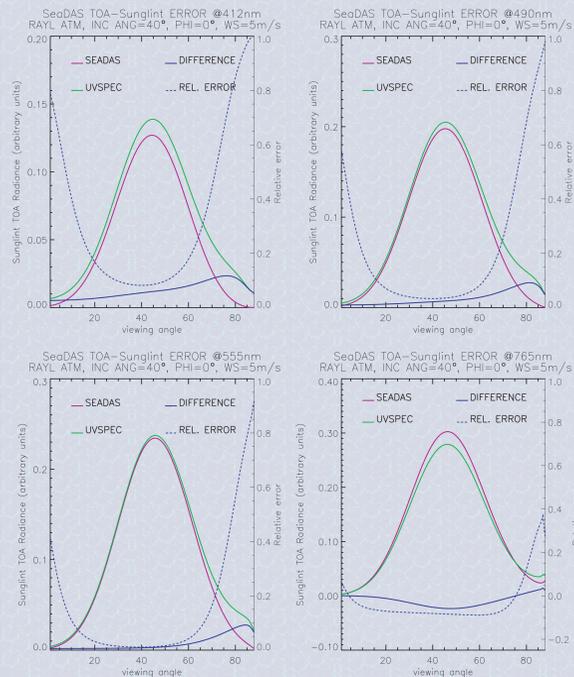
## Sunlight correction in

### SeaDAS

Currently the SeaDAS sunlight correction is:

- ~ based on 1D Cox/Munk (no wind direction)
- ~ not attempted if the glint is too strong
- ~ using a simple direct transmittance calculation to determine the glint's contribution at the TOA

We are investigating a more sophisticated method that ~ uses accurate RT calculations to determine the glint's contribution at the TOA ~ will enable retrievals further into the glint region, reducing amount of wasted data ~ is easy to implement



Above: Sunlight TOA calculated using UVSPEC and SeaDAS at 4 different SeaWiFS channels. The green curve is the difference between the two models.

## Measuring Sunlight

There are two fundamental ways of looking at sunlight:

- time averaged (Cox/Munk) - the regime for low spatial resolution, long sampling (exposure) times
- instantaneous (time resolved) - hyper spatial, fast sampling rates (short exposure)

Time averaged (BRDF) view of sunlight is needed for / can be used for:

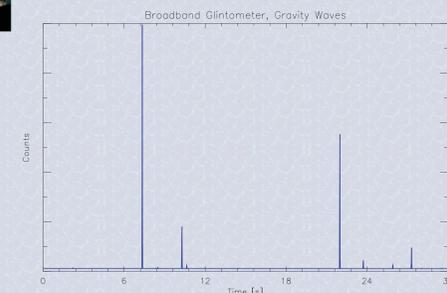
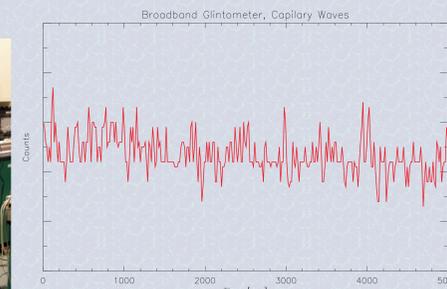
- detecting departure from Gaussian statistics
- input to RT models (forward models)

A dynamic view of sunlight is needed for / can be used for:

- determination of types and magnitude of shear forces (currents, wind)



Above: Pictures from our lab showing the CIMEL sun photometer (left) and our own "glintometer". The CIMEL is capable of measuring radiances at 6 different wavelengths with polarization. Limited sampling rates prevents dynamic (time-resolved) glint measurements. Our glintometer can sample from DC to a maximum rate of 3.5kHz, enough to resolve the evolution of even the fastest "glint atoms". The glintometer measures broad-band radiances, but we plan to include filters in future versions of the instrument.



Above: Examples of the dynamic nature of glint measured by our glintometer. Top: small, wind-driven capillary waves. Bottom: large gravity waves.

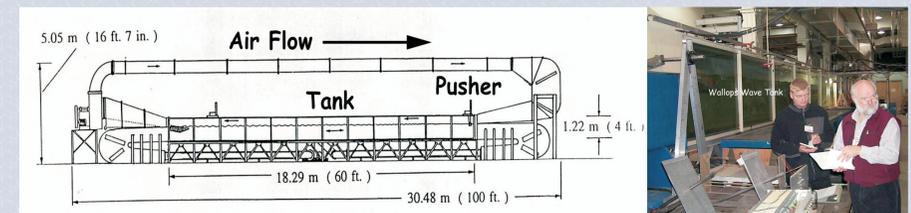
## Future Modeling and Measurement Plans

Development of complete Monte Carlo model with realistic surface realizations. This model will include

- shadowing effects, re-emerging photons, scattering by bubbles and foam
- polarization effects

On the experimental side we plan to utilize the Wallops Island wave tank facility to

- check the performance of current models
- investigate the effects of bubbles/foam, wind/current, thermal stability
- investigate the polarization features of the glint



Above: The NASA Wallops wave tank facility.

## SUMMARY

Our sunlight studies are aimed at

- improving current methods for sunlight corrections in ocean color retrievals
- developing methods for determining atmospheric parameters using sunlight as a tool
- performing measurements of the dynamic, as well as time-averaged, sunlight
- measuring and modeling of polarized sunlight